# THERMAL & STATISTICAL PHYSICS SSP3133

# **QUANTUM STATISTICS**

### DR WAN NURULHUDA WAN SHAMSURI

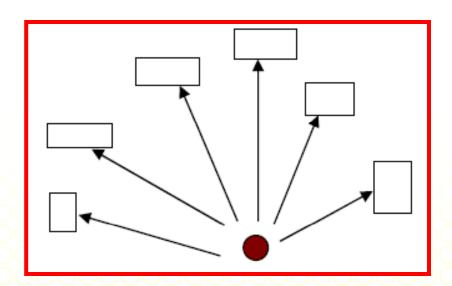
Acknowledgement: PROFESSOR DR RAMLI ABU HASSAN

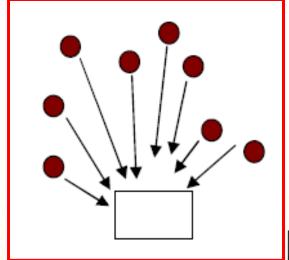




# Classical & quantum statistics

Classical statistics: focus on individual paticle, which could occupy any of several possible quantum states Quantum statistics: focus on individual quantum state, which could be occupied by various particles





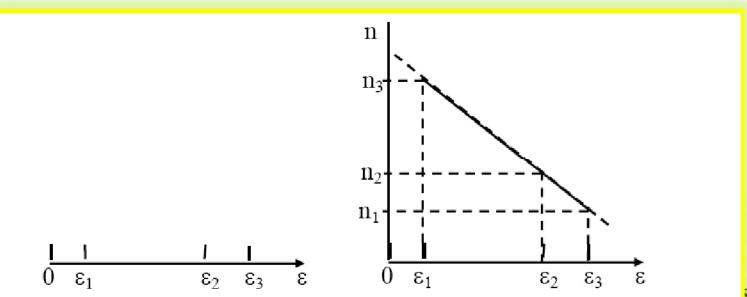




Statistics of systems obeying quantum physics: the statistical description of systems of particles that are subject to the laws of quantum physics rather than classical physics

### Problems in QM

- i. Finding the spectrum of the accessible states
- ii. Finding the average number of particles occupying each state







# the occupation number: n

Note: consider the system to a certain quantum state s. if  $\varepsilon_s$ : the energy of single particle in this state the probability that there are n particles in this state

Where

$$P_{S} = Ce^{-n\beta(\varepsilon_{S}-\mu)}$$

$$C = \left(\sum_{n} e^{-n\beta(\varepsilon_{S} - \mu)}\right)^{-1}$$



The average number of particles occupying the state

$$\overline{n} = \sum_{n} Pn = \left[\sum_{n} e^{-n\beta(\varepsilon_{S} - \mu)}\right]^{-1} \sum_{n} e^{-n\beta(\varepsilon_{S} - \mu)} n$$

let 
$$x = \beta(\epsilon_S - \mu)$$

$$\overline{n} = \frac{\sum_{n} e^{-nx} n}{\sum_{n} e^{-nx}} = \frac{-\frac{\partial}{\partial x} \sum_{n} e^{-nx}}{\sum_{n} e^{-nx}} = -\frac{\partial}{\partial x} \ln \left( \sum_{n} e^{-nx} \right)$$
(\*)

# TWO types of particles:

# FERMIONS & BOSONS

FERMIONS: Only one particle may occupy a given state

BOSONS: no restriction on the number of particles that may occupy a given state





Note:

# 2 particles in 3 energy states

i) MB distribution – 9 possible configurations

	T		
Configuration	State 1	State 2	State 3
1	AB	-	-
2	-	AB	-
3	-	-	AB
4	A	В	-
5	В	A	-
6	A	-	В
7	В	-	A
8	-	A	В
9	-	В	A



ii) BE distribution – 6 possible configurations

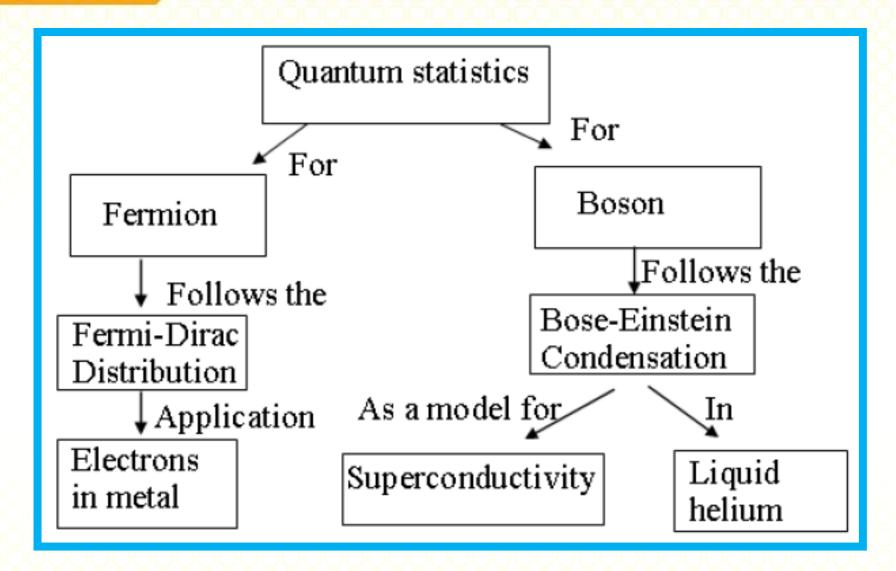
Configuration	State 1	State 2	State 3
1	AA	-	-
2	-	AA	-
3	-	-	AA
4	A	A	-
5	A		A
6		A	A



iii)

FD distribution – 3 possible configurations

Configuration	State 1	State 2	State 3
1	A	A	
2	A	_	A
3	-	A	A



# **Bose - Einstein statistics**

- -Indian physicist Satyendra Bose derived the quantum statistics of photons by assuming them as indistinguishable particles.
- -Albert Einstein applied it to atoms.

-Note: **atoms:** the particle number is conserved (if they are trapped in a box or a magnetic field)

**Photons** can randomly be emitted and absorbed (e.g. from the walls of a box).



B-ED - Class of statistics that applies to elementary particles called **bosons**, which include the photon, pion, and the W and Z particles.

Particles behaving according to the B-E distribution are called **bosons**--- having spin values 0, 1, 2,..etc.

Q-statistics --- Bose-Einstein statistics

For B-E stat.

$$\sum_{n} e^{-nx} = \frac{1}{1 - e^{-x}}$$
 FOR X>0



# From (\*)

$$\overline{n}_{bosons} = -\frac{\partial}{\partial x} \ln \left( 1 - e^{-x} \right)^{-1} = \frac{\partial}{\partial x} \ln \left( 1 - e^{-x} \right)^{-1}$$

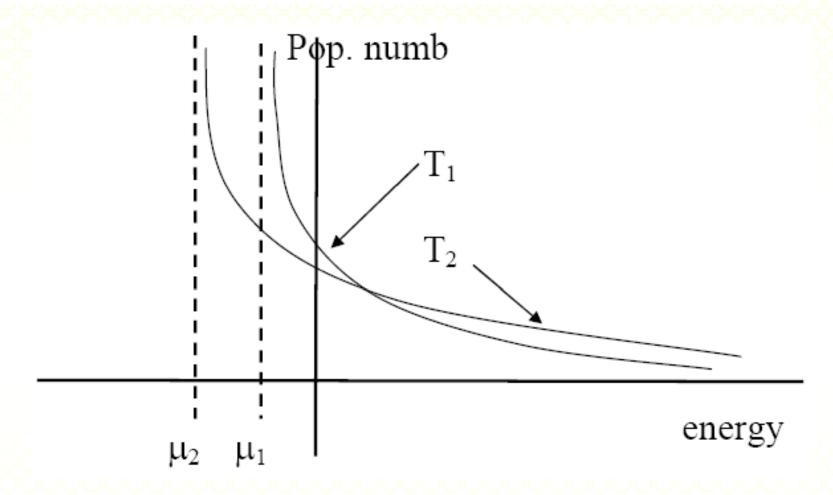
$$= \frac{e^{-x}}{1 - e^{-x}} = \frac{1}{e^x - 1}$$

But 
$$x = \beta(\epsilon_S - \mu) > 0$$

$$\overline{n} = \frac{1}{e^{\beta(\varepsilon - \mu)} - 1}$$



Note: at  $\varepsilon = \mu$ , the population number would be infinite



At low temperatures, an unlimited number of bosons can be in the same energy state, a phenomenon called condensation.

$$f(E) = \frac{1}{Ae^{E/kT} - 1}$$

f(E): the probability a particle will have energy E

A: for photon equals to one

e<sup>E/kT</sup>: the exponential dependence on energy and temperature

-1: the quantum difference which arises from the fact the particles are indintinguishable





$$f(E) = \frac{1}{e^{\frac{E-\mu}{kT}} - 1}$$

# where:

E is the energy  $k_B$  is Boltzmann's constant T is absolute temperature  $\mu$  is the chemical potential

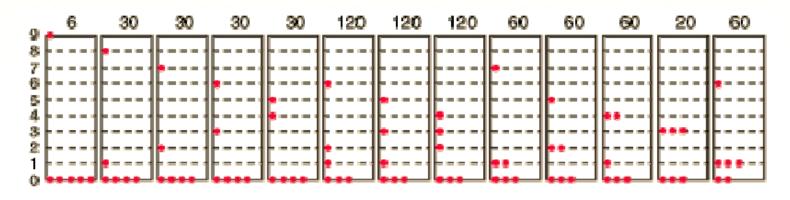


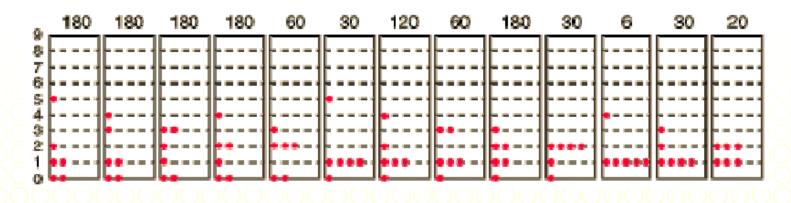


# Example

How many ways can you distribute 9 units of energy among 6 identical, distinguishable / indistinguishable bosons?

Note: for M-B stat

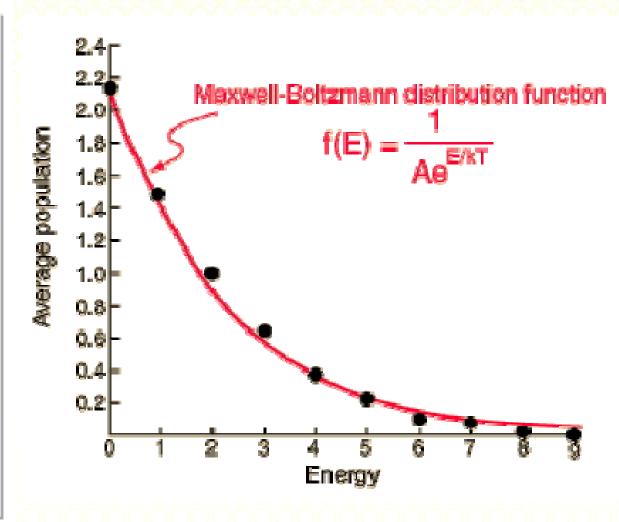








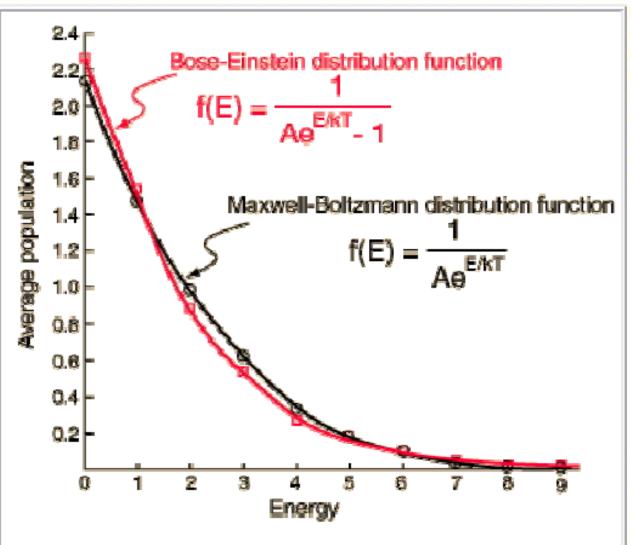
Energy level	Average number
0	2.143
1	1.484
2	0.989
3	0.629
4	0.378
5	0.210
6	0.105
7	0.045
8	0.015
9	0.003







Energy level	Average number Maxwell - Boltzma nn	Average number Bose- Einstein
0	2.143	2.269
1	1.484	1.538
2	0.989	0.885
3	0.629	0.538
4	0.378	0.269
5	0.210	0.192
6	0.105	0.115
7	0.045	0.077
8	0.015	0.038
9	0.003	0.038







# Fermi-Dirac statistics

- Class of statistics that applies to particles called fermions.
- Fermions have half-integral values of the quantum mechanical property called spin and that two fermions cannot exist in the same state.
- Protons, neutrons, electrons, and many other elementary particles are fermions.

The Fermi-Dirac function f(E) gives the probability that an electron (for example) has energy E at temperature T given by:

$$f(E) = \frac{1}{e^{\frac{E - E_f}{kT}} + 1}$$

Where  $E_f$  is the Fermi energy.



The **Fermi energy** ( $E_f$ ) of a system of non-interacting fermions is the smallest possible increase in the ground state energy when exactly one particle is added to the system

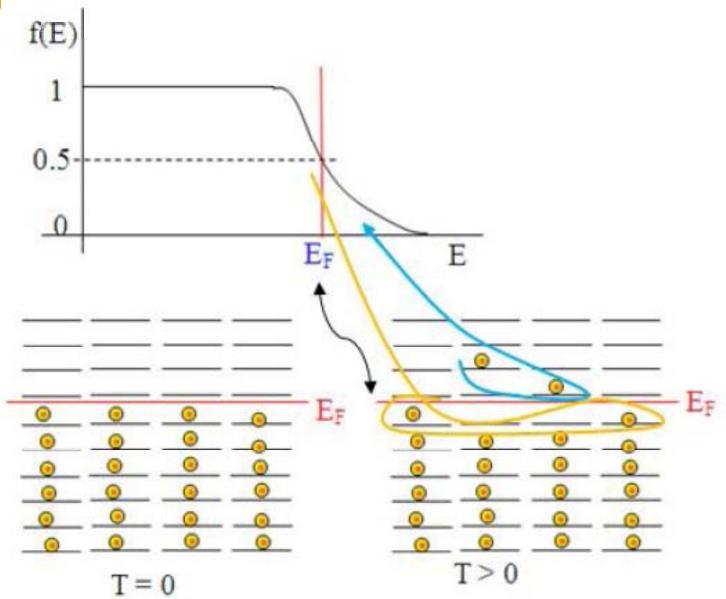
Is equivalent to the chemical potential ( $E_f = \mu$ ) of the system in its ground state at T=0 K.

The **Fermi level** is the top of the collection of electron energy levels at absolute zero temperature.

The Fermi energy or Fermi level,  $E_f$  is defined as the energy at which the probability of occupation of the electron state is 1/2.









# The Fermi-Dirac Distribution

$$f(E) = \frac{1}{\frac{E - E_f}{e^{kT} + 1}}$$

At 0K fermions will fill up all available energy states below a level  $E_F$  with one particle.

At higher temperature some are elevated to levels above E<sub>F</sub>

Term + 1: the quantum difference which arises from the fact that the particles are indistinguishable



Fermions are particles with half-integer spin.

By the Pauli Exclusion Principle, only one fermions can occupy a given state at a time.

That means that the occupation number for state s can be only  $n_s = 0$  or  $n_s = 1$ 

-conservation of particle number. 
$$\sum_{s} n_{s} = N$$

Note: 
$$\sum_{n=0}^{1} e^{-nx} = 1 + e^{-x}$$
 n can be 1 or 0





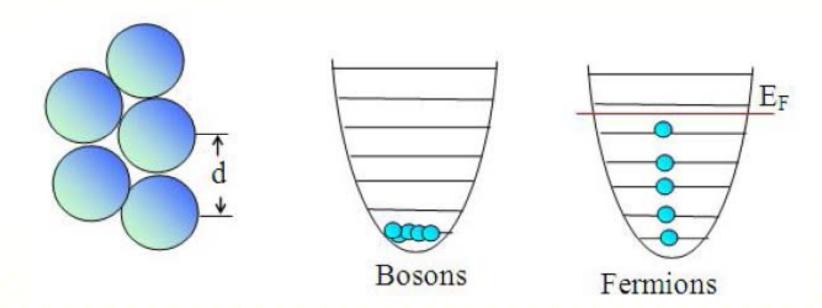
The mean number of particles in state s

$$\begin{split} \overline{n}_s &= -\frac{1}{\beta} \frac{\partial \ln Z}{\partial \varepsilon_s} \\ &= -\frac{1}{\beta} \frac{\partial}{\partial \varepsilon_s} \Big[ \ln(1 + e^{-\beta(\varepsilon_s - \mu)} + \dots \Big] \\ &= \frac{e^{-\beta(\varepsilon_s - \mu)}}{1 + e^{-\beta(\varepsilon_s - \mu)}} \end{split}$$



And finally 
$$\overline{n}_s = \frac{1}{e^{\beta(\varepsilon_s - \mu)} + 1}$$

This is known as the <u>Fermi-Dirac Distribution</u> also called <u>Fermi function</u>



at T = 0, Bosons: the energy is  $\mu$  Fermions: the highest energy is  $E_F$ 



### Fermi-Dirac distribution / Fermi Function

The Fermi function f(E) gives the probability that a given available electron energy state will be occupied at a given temperature.

Note: 
$$\mu = \left(\frac{\partial E}{\partial N}\right)_{S,V} = -T\left(\frac{\partial S}{\partial N}\right)_{E,V}$$
  $\mu \equiv -kT\alpha$ 

The Fermi energy is defined as  $E_F = -\alpha(T)kT$ 

 $\alpha(T)$ : the number of particles in a particular system at temperature T,





$$f(E) = \frac{1}{e^{\frac{E - E_f}{kT}} + 1}$$

At 0 K no electrons can be above the valence band, since none have energy above the Fermi level and there are no available energy states in the band gap.

At 
$$T = 0$$
 K,  $\beta = 1/kT = \infty$ 

If 
$$E \le E_f$$

If  $E > E_f$ 

Then 
$$f(E) = 1$$

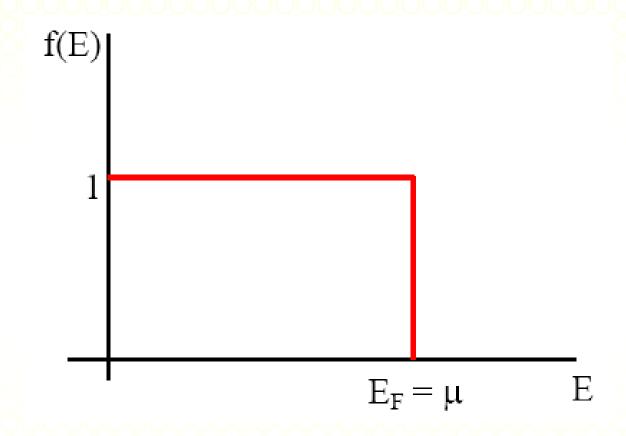
Then 
$$f(E) = 0$$

$$e^{\frac{E - E_f}{kT}} = e^{-\infty} = 0$$

$$e^{\frac{E-E_f}{kT}} = e^{\infty} = \infty$$











Increasing T

If 
$$E = E_f$$
 then  $f(E) = \frac{1}{2}$ 

If 
$$E > EF$$
, let say  $E \ge E_F + 3kT$ 

Then 
$$e^{\frac{E-E_f}{kT}} = e^3 > 1$$

Approximation:

$$f(E) = e^{\frac{-(E - E_f)}{kT}}$$

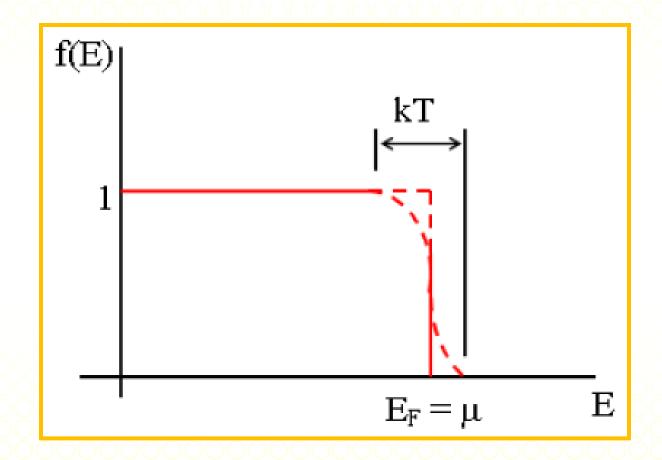
 $If \: E <\!\! E_F \quad \ \ let \: say \: E \leq E_F \text{ - } 3kT$ 

Then 
$$e^{\frac{E-E_f}{kT}} = e^{-3} << 1$$



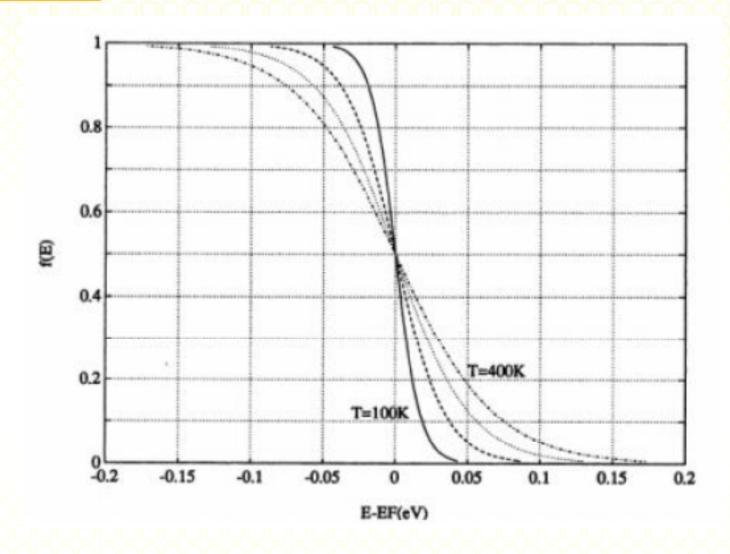
Approximation:

$$f(E) = 1 - e^{\frac{(E - E_f)}{kT}}$$

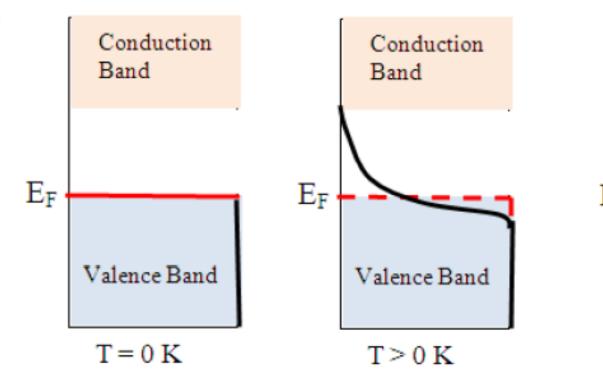


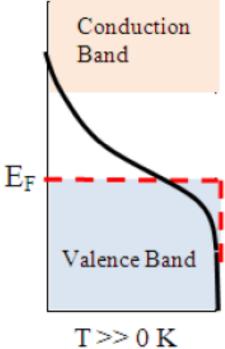






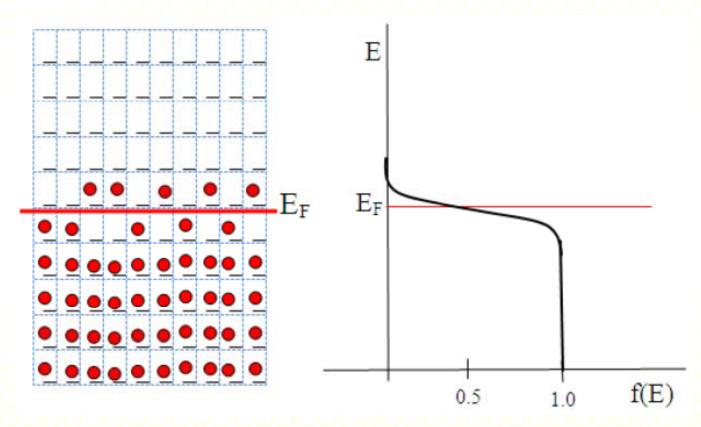






At high temperature, some electrons can reach the conduction band and contribute to electric current.

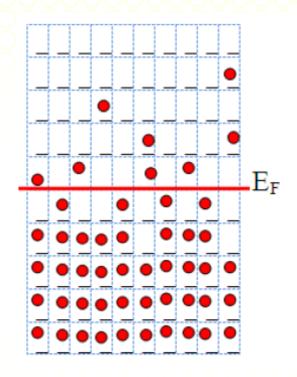
### T~300 K

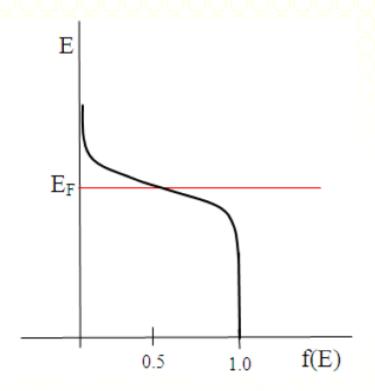






### T~ 800 K

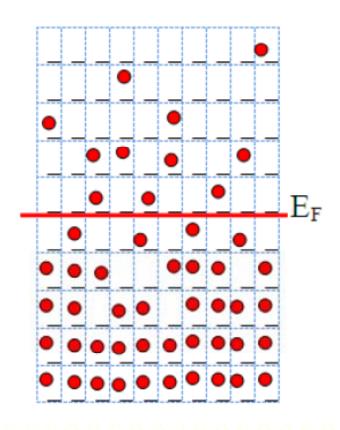


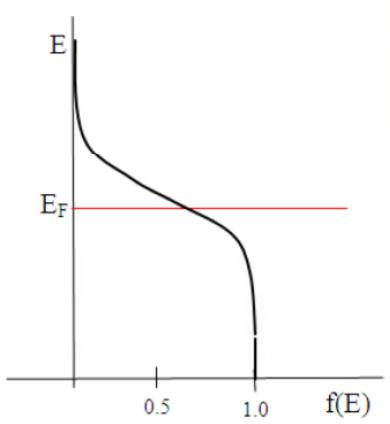






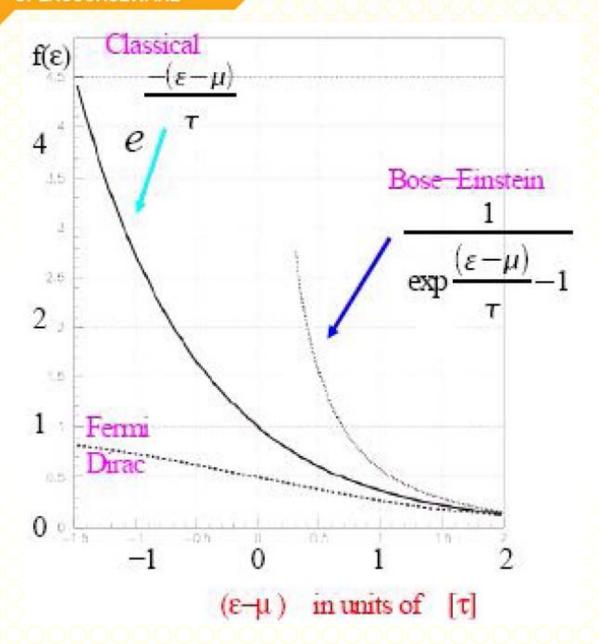
### T~1300 K















$$f(\varepsilon)_{BE} = \frac{1}{\exp \frac{(\varepsilon - \mu)}{\tau} - 1}$$

$$f(\varepsilon)_{FD} = \frac{1}{\exp \frac{(\varepsilon - \mu)}{\tau} + 1}$$

Large  $(\varepsilon - \mu)/T$  BE & FD approach classical distribution, MB

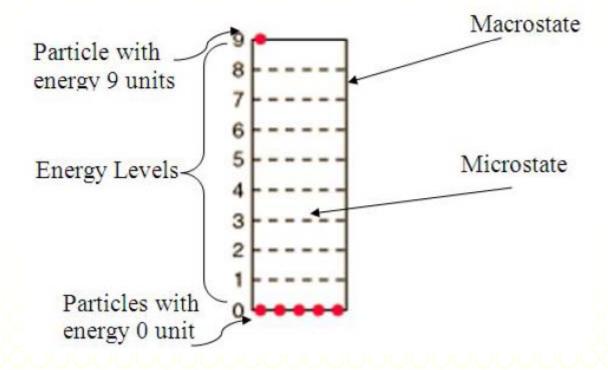
$$f(\varepsilon)_{classic} = e^{-\frac{(\varepsilon-\mu)}{\tau}}$$



## Example

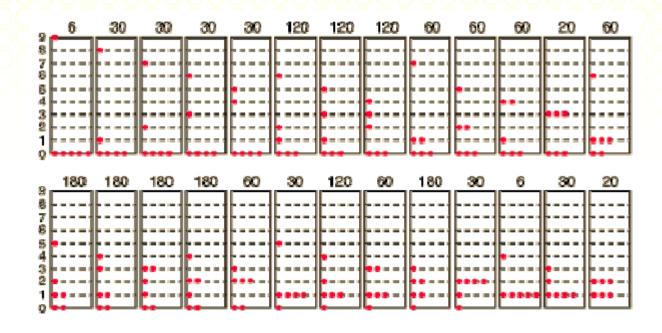
The distribution of 9 units of energy among 6 identical particles

## Note:





# 1. MBD (each particle is presumed to be distinguishable)



-The total number of different states for distinguishable particles is 2002 = 6+30+30+...







No. of microstates = 
$$\frac{N!}{n_1!n_2!n_3!...}$$

N: total number of particles

n<sub>i</sub>: number of particles in level i

The first macrostates there are 6 microstates

$$\frac{6!}{5!1!} = 6$$

The second macrostates there are 30 microstates

$$\frac{6!}{4!1!1!} = 30$$
 Note:  $0! = 1$ 





The number of microstates for ε units of energy among N equally probable states

$$\Omega(N,\varepsilon) = \frac{(\varepsilon + N - 1)!}{\varepsilon!(N - 1)!}$$

For this case becomes

$$\Omega(N,\varepsilon) = \frac{(9+6-1)!}{9!(6-1)!} = \frac{14.13.12.11.10}{5.4.3.2.1} = 2002$$



To find average particles per energy level

For each macrostate, the number of particles in a given energy level is multiplied by the number of microstates.

The sum of those products is divided by the total number of microstates

$$n_j = \sum_i n_{ij} P_i$$

 $n_{ij}$  = number of particles of energy  $E_j$  in microstate i

P<sub>i</sub>= number of microstates in macrostate I divided by the total number of microstates

 $n_j$  = average number of particles in energy level j





the average number of particles in energy level 0

$$n_0 = 5x \frac{6}{2002} + 4x \frac{30}{2002} + 4x \frac{30}{2002} + 4x \frac{30}{2002} + 4x \frac{30}{2002} + \dots$$

$$= \frac{1}{2002} \{5x6 + 4(4x30) + 3(3x120) + 3(3x60) + (3x20) + (2x60)\}$$

$$+4(2x180)+(2x60)+(1x30)+(1x120)+(1x60)+(1x180)+(1x30)$$

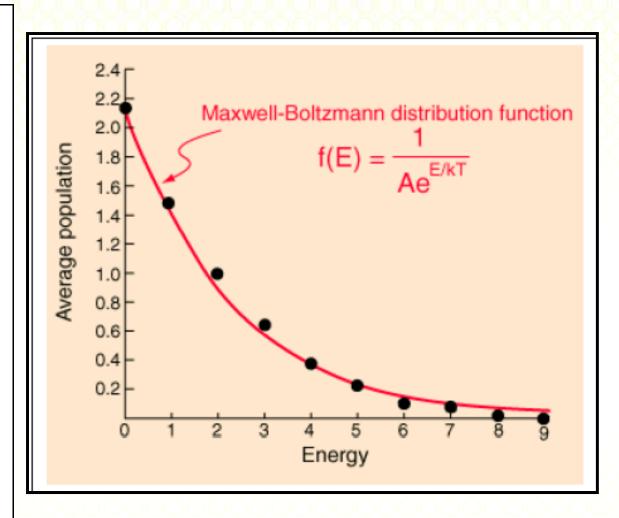
$$n_0 = 2.143$$

continue to calculate  $n_1, n_2, n_3, \ldots, n_9$ 





Energy level	Average number	
0	2.143	
1	1.484	
2	0.989	
3	0.629	
4	0.378	
5	0.210	
6	0.105	
7	0.045	
8	0.015	
9	0.003	





# 2. BED (identical, indistinguishable bosons)

there are 26 microstates arising from 26 macrostates

The average occupancy is the sum of the numbers of particles in a given energy state over all the 26 distributions divided by 26.

For average particles in energy level 0,

$$n_0 = (5+4+4+....)/26 = 59/26 = 2.269$$

continue ....

$$n_1 = \ldots, \qquad n_2 = \ldots \qquad \ldots$$





# 3. FDD (fermions must obey the Pauli exclusion principle).

if the particles are electrons,

- a maximum of two particles can occupy each spatial state since there are two spin states each.
- -26 possible configurations for distinguishable particles are reduced to the 5 states which have no more than two particles in each state.





The average occupancy per energy level

the sum of the numbers of particles in a given energy state over all the 5 distributions divided by 5.

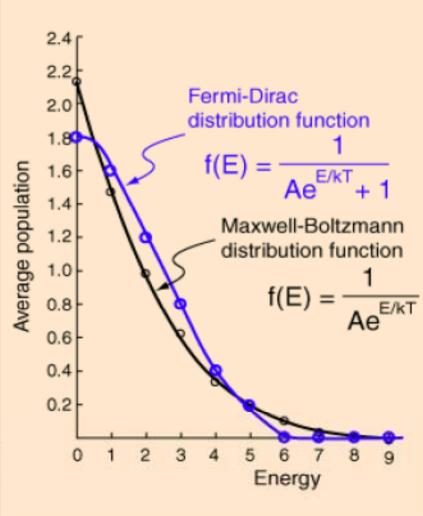
$$n_0 = 9/5 = 1.8$$

$$\dots n_1, n_2, n_3, \dots$$



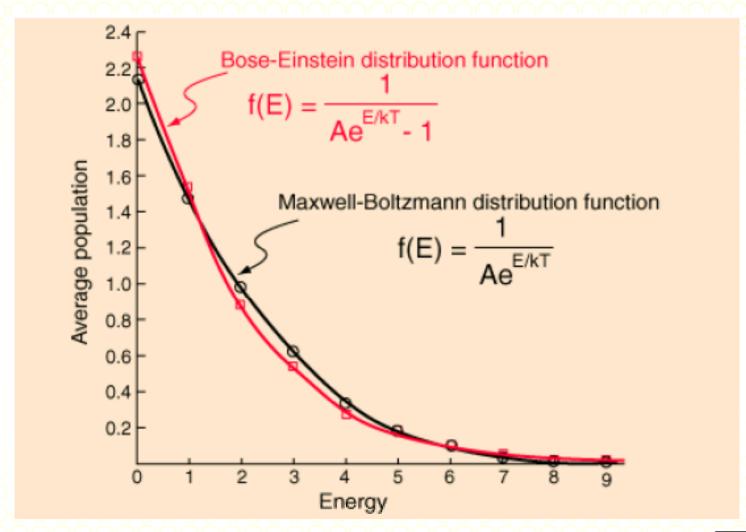


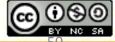
Е	Average number Maxwell- Boltzmann	Average number Bose- Einstein	Average number Fermi- Dirac
0	2.143	2.269	1.8
1	1.484	1.538	1.6
2	0.989	0.885	1.2
3	0.629	0.538	0.8
4	0.378	0.269	0.4
5	0.210	0.192	0.2
6	0.105	0.115	0
7	0.045	0.077	0
8	0.015	0.038	0
9	0.003	0.038	0













# PHEW...!!



## REFERENCES:

- 1. REAF, F: "Fundamentals Of Statistical And Thermal Physics", McGraw-Hill.
- 2. KITTEL & KROMER: "Thermal Physics", W.H. Freeman & Company.

